

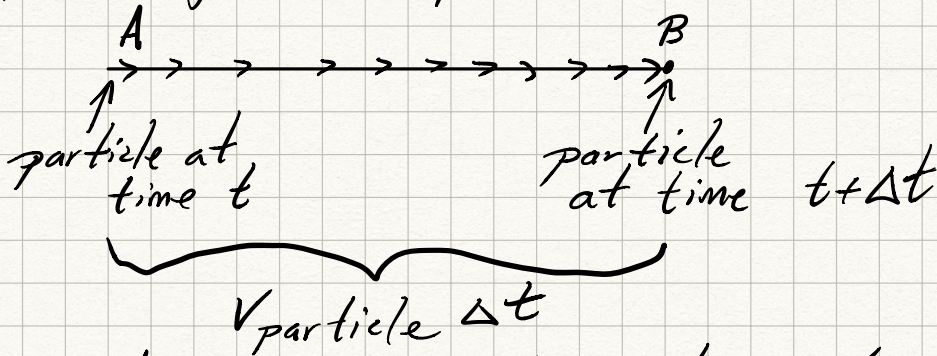
Third SR Homework - Solutions

Problems 3-8, 3-9 and 3-10

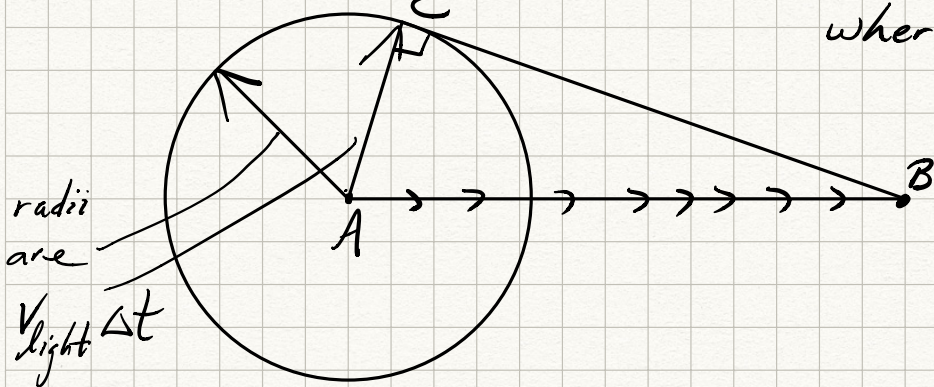
Problem 3-8 Čerenkov Radiation

(a)

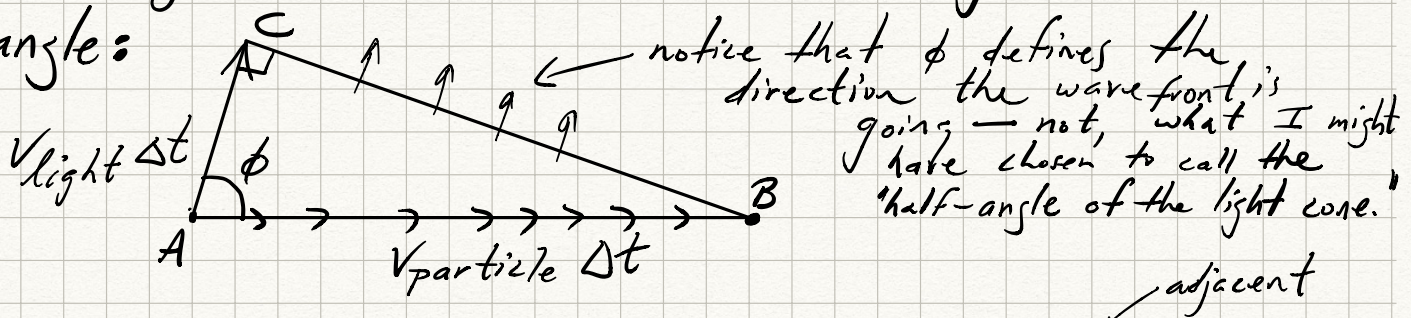
First, draw just the particle motion without the light:



Next, add a circle with radius $v_{\text{light}} \Delta t$ representing where the light emitted at time t has gotten to at time $t + \Delta t$



Then add the tangent to the circle BC and a radius (that is perpendicular to the tangent) AC. Then focus on just the triangle:



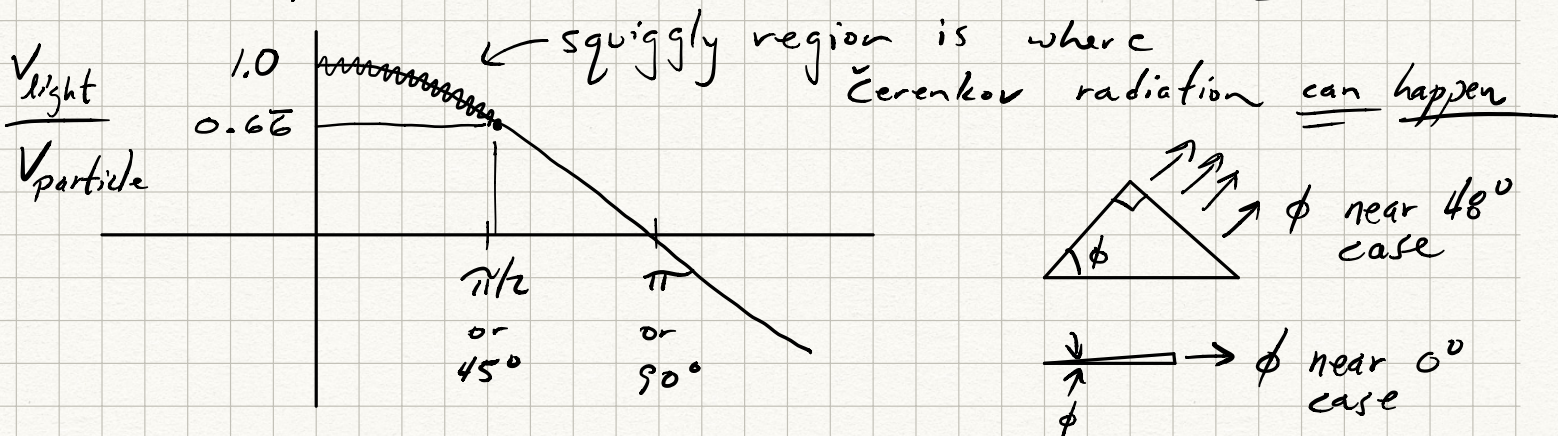
Since it is a right triangle with $A = v_{\text{light}} \Delta t$ and $H = v_{\text{particle}} \Delta t$, we have

$$\cos \phi = \frac{A}{H} = \frac{v_{\text{light}} \Delta t}{v_{\text{particle}} \Delta t} = \frac{v_{\text{light}}}{v_{\text{particle}}}$$

(b) In Lucite, $v_{light} = \frac{2}{3}c = \frac{2}{3} \cdot 3 \times 10^8 \frac{m}{s} = 2 \times 10^8 \frac{m}{s}$

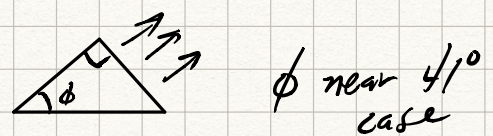
To have Čerenkov radiation, the particle must be going faster than v_{light} . The absolute maximum speed that any particle can go is (anything just less than $c = 3 \times 10^8 \frac{m}{s}$). Using the limiting case of c itself (even though it can never be attained) for $v_{particle}$, we have

$$\cos \phi = \frac{v_{light}}{c} = \frac{2}{3} \quad \phi = \cos^{-1} \frac{2}{3} = 48^\circ$$



(c) In water $v_{light} = 0.75c = \frac{3}{4}c = \cos \phi$

$$\phi = \cos^{-1} \frac{3}{4} = 41^\circ$$



3-9 Aberration of Starlight

(a) It seems we should analyze in sun frame to get ψ ?



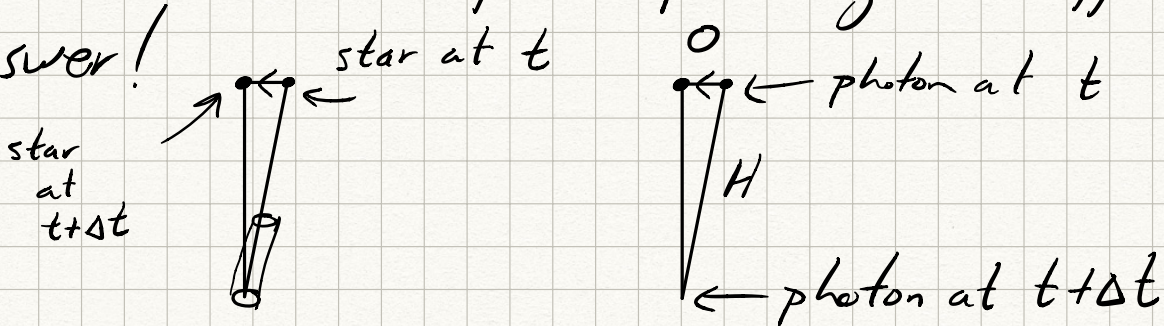
$$3-9(a) \text{ (CONT'D)} \quad O = v_{\text{earth}} \Delta t \quad A = v_{\text{photon}} \Delta t$$

$$\tan \psi = \frac{O}{A} = \frac{v_{\text{earth}} \Delta t}{c \Delta t}$$

$$= \frac{v_{\text{earth}}}{c} = v_{\text{earth}} \leftarrow \text{in units where } c=1$$

We have found the direction the telescope must point — in the Sun's frame!

In the Earth's frame, we get a different answer!



$$O = v_{\text{earth}} \Delta t$$

$$H = c \Delta t$$

$$\sin \psi = \frac{O}{H} = \frac{v_{\text{earth}} \Delta t}{c \Delta t} = \frac{v_{\text{earth}}}{c} = v_{\text{earth}}$$

\uparrow
in units where $c=1$

Don't we want to know the angle of the telescope in our frame!

So the second formula is the one we want!

(b) Plug in numbers

$$\psi = \sin^{-1} v_{\text{earth}}$$

$$= \sin^{-1} 10^{-4}$$

$$v_{\text{earth}} = \frac{30 \text{ km/s}}{3 \times 10^8 \text{ m/s}}$$

$$= 10^{-4}$$

The sin of an extremely small angle is just the angle in radians! So

$$\psi = 10^{-4} \text{ radians}$$

3-9(b) (CONT'D) multiply by $\frac{180}{\pi}$

Convert to degrees, get $\psi = 0.0057$ degrees

Convert to minutes of arc $\psi = 0.34$ minutes of arc

Convert to seconds of arc $\psi = 21$ seconds of arc ← multiply by 60

(c) Compare Sun-frame formula with
with Earth-frame formula ← naive, not the angle we measure

$$\psi = \tan^{-1} 0.001 = 0.005729577932^\circ \leftarrow \text{my calculator displays 10 sig figs}$$

$$\psi = \sin^{-1} 0.001 = 0.005729577960^\circ \leftarrow \text{only last two differ}$$

The absolute best angular measurement are made by the Gaia satellite which is the successor to the Hipparcos satellite. It can do angles between stars to about 10^{-11} radians. That's 0.000000000600° . Not good enough for two reasons: (1) it's 20x as big as the effect we are looking for, and (2) that's its best measurement of the angles between stars. As far as I know, it cannot measure absolute angles with this precision.

(d) Re-evaluate with $v_{\text{rocket}} = 0.5$

$$\psi = \tan^{-1} 0.5 = 26.56^\circ \leftarrow \text{Sun-frame}$$

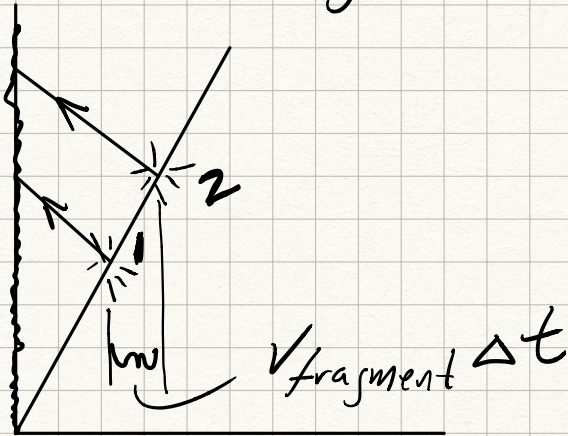
$$\psi = \sin^{-1} 0.5 = 30^\circ \leftarrow \text{Rocket frame}$$

3-10 (a) Expanding universe

$\Delta\tau$ = proper time between flashes (time in frame of emitter)

$\Delta t_{\text{reception}}$ = time between flashes (as observed by receiver)

Space-time diagram in receiver's frame:



squiggly line is world-line of receiver

$$\Delta t = \frac{1}{\sqrt{1 - v_{\text{fragment}}^2}} \Delta\tau$$

The second photon has to go $v_{\text{fragment}} \Delta t$ further.

we don't have to re-derive this.

So it is received not just Δt later, but $\Delta t + v_{\text{fragment}} \Delta t$ later

Now rewrite this in terms of $\Delta\tau$

$$\begin{aligned} \Delta t_{\text{reception}} &= (1 + v_{\text{fragment}}) \frac{1}{\sqrt{1 - v_{\text{fragment}}^2}} \Delta\tau \\ &= (1 + v_{\text{fragment}}) \frac{1}{\sqrt{(1 + v_{\text{fragment}})(1 - v_{\text{fragment}})}} \Delta\tau \quad \text{factor} \\ &= \sqrt{\frac{1 + v_{\text{fragment}}}{1 - v_{\text{fragment}}}} \Delta\tau \end{aligned}$$

I will discuss the discussion part of 3-9(a) in class.

3-9(b) I will also discuss this discussion question in class

(c) Quasars have been seen with $\frac{\Delta t_{\text{reception}}}{\Delta \tau} = 5.9$
What is v_{quasar} ?

Call the ratio y . Solve

$$y = \sqrt{\frac{1 + v_{\text{fragment}}}{1 - v_{\text{fragment}}}} \quad \text{for } v_{\text{fragment}} \text{ by first squaring each side.}$$

$$y^2 = \frac{1 + v_{\text{fragment}}}{1 - v_{\text{fragment}}} \Rightarrow y^2 (1 - v_{\text{fragment}}) = 1 + v_{\text{fragment}}$$

$$\Rightarrow y^2 - 1 = (y^2 + 1) v_{\text{fragment}}$$

$$\Rightarrow v_{\text{fragment}} = \frac{y^2 - 1}{y^2 + 1} = 0.944$$

Almost 95% the speed of light.